

# §. Lagrangian EOM with constraint(s)

Let us consider one constraint only.  
(generalization to multiple constraints straight-f.)

Assume ~~that~~  $\sum_k a_k \dot{q}_k + a_t = 0$ .

a constraint that satisfies  $\rightarrow$  functions of  $q_k, t$

A special commonplace sub-class of all possible constraints  
 $F(q_k, \dot{q}_k, t) = 0$   
all  $k$ 's  $1 \rightarrow n$

In general, this cannot be integrated (non-holonomic)  
If it can be integrated, then we have a holonomic constraint.  
(e.g.  $v = R\omega$  for a rolling w/o ~~slipping~~ <sup>slipping</sup>)

Holonomic constraint:  
 $F(q_k, \dot{q}_k, t) = G(q_k, t) = 0$

The above equation can be written as  
 $\sum_k a_k d\dot{q}_k + a_t dt = 0$

For a virtual displacement  $\delta t = 0$   
 $\sum_k a_k \delta \dot{q}_k = 0 \dots (*)$

↑ simple.  
Can reduce the # of generalized coordinates by 1 per constraint.

$\int_1^2 dt \sum_k \left( \frac{\partial L}{\partial \dot{q}_k} - \frac{d}{dt} \frac{\partial L}{\partial q_k} \right) \delta q_k = 0$   
Add 0 times  $\lambda(t)$  --- Lagrange multiplier <sup>undetermined</sup>

$$\int_1^2 dt \sum_k \left[ \frac{\partial L}{\partial \dot{q}_k} - \frac{d}{dt} \frac{\partial L}{\partial q_k} + \lambda a_k \right] \delta q_k = 0$$

Not all  $q_k$ 's are independent!

But only one constraint.

From  $k=1, \dots, n$ , if we ~~remove~~ exclude  
( $n = \text{D.O.F.}$ )

one of  $q_k$ 's, then the rest are independent!

So, consider  $q_2, \dots, q_n$  as indep. vars.

As for  $q_1$ , choose  $\lambda(t)$  so that

$$\frac{\partial L}{\partial q_1} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_1} + \lambda(t) a_1 = 0$$

As for other  $q_k$ 's, ( $k=2, \dots, n$ ) the variational principle as before

$$\Rightarrow \frac{\partial L}{\partial q_k} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_k} + \lambda(t) a_k = 0$$

Result? This equation is valid for all  $k$ 's including  $k=1$ .

Price?  $n+1$  unknowns  $q_k, \lambda(t)$   
 $k=1, \dots, n$

only  $n$  Lagrange eq's?! No worries. We have the constraint equation

$$\sum_k a_k \dot{q}_k + a_t = 0$$

Meaning of  $\lambda(t)$ : generalized force of constraint

$$-\frac{\partial L}{\partial q_1} + \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_1} = + \lambda(t) a_1$$

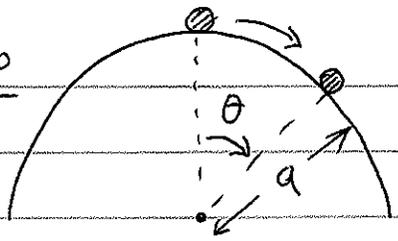
Lagrangian mechanics is cool/sneaky in this regard. No need to know the constraint force to start with. It is an outcome! (friction... normal force...)

Why write in this form??

See L9-⑦!  
 $-\frac{\partial L}{\partial q_j} + \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_j}$   
=  $Q_j$   
generalized force

$\theta = \dot{\theta} = 0$  at  $t=0$  L 11 - (3)

★  
Ex. 7.10



$r, \theta$   
 $T = \frac{1}{2} m \{ \dot{r}^2 + (r\dot{\theta})^2 \}$

$U = mgr \cos \theta$

$L = \frac{1}{2} m \{ \dot{r}^2 + (r\dot{\theta})^2 \} - mgr \cos \theta$

①  $\frac{d}{dt} \frac{\partial L}{\partial \dot{r}} - \frac{\partial L}{\partial r} = \lambda \left( \frac{\partial f}{\partial r} \right)^1$   $f(r, \theta) = r - a = 0$

②  $\frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}} - \frac{\partial L}{\partial \theta} = \lambda \left( \frac{\partial f}{\partial \theta} \right)_0$   $\frac{\partial f}{\partial r} = 1$   
 $\frac{\partial f}{\partial \theta} = 0$

①  $m\ddot{r} - mr\dot{\theta}^2 + mg \cos \theta = \lambda$

②  $\frac{d}{dt} (mr^2\dot{\theta}) = mgr \sin \theta$   $\dot{\theta} \frac{d\dot{\theta}}{d\theta} = \frac{\dot{\theta}^2}{2}$   $\leftarrow$  cons. E

③  $r = a$  ②  $\Rightarrow \ddot{\theta} = \frac{g}{r} \sin \theta$

①  $\Rightarrow -mr\dot{\theta}^2 + mg \cos \theta = \lambda$

④  $T + U = E = \frac{1}{2} m \{ \dot{r}^2 + (r\dot{\theta})^2 \} + mgr \cos \theta = mga$

$\lambda$ : normal force  $\left| \frac{1}{2} a \dot{\theta}^2 + g \cos \theta = g \right|$

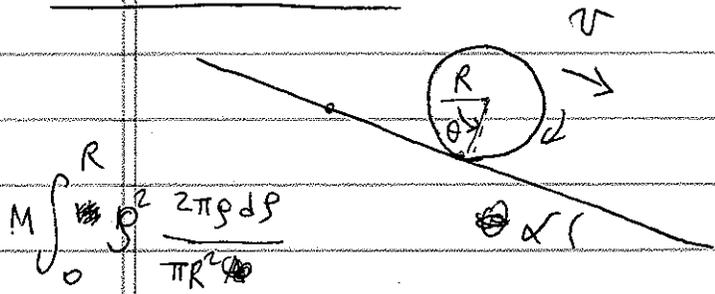
$\lambda = 0 \dots a \dot{\theta}^2 = g \cos \theta$   
 $\lambda > 0 \dots$  ok.  
 $\lambda < 0 \dots$  not ok.  $\frac{3}{2} g \cos \theta = g \therefore \cos \theta = \frac{2}{3}$

The particle takes off from the surface.  $\theta = \cos^{-1} \frac{2}{3}$

Was that easier than doing  $\left( \begin{array}{l} mg \cos \theta = m \frac{v^2}{r} \\ \frac{1}{2} mv^2 + mg a \cos \theta = mga \end{array} \right)$  ?  
No!

do it

Ex 7.9



Constraint  $v = R \dot{\theta}$

$s = R\theta$

$s$ : distance traveled from rest

$$M \int_0^R \rho^2 \frac{2\pi \rho d\rho}{\pi R^2}$$

$$T = \frac{1}{2} M \dot{s}^2 + \frac{1}{2} \cdot \frac{1}{2} M R^2 \dot{\theta}^2 \quad U = -Mg s \sin \alpha$$

$v = \dot{s}$

$$= \frac{1}{2} M \dot{s}^2 + \frac{1}{4} M R^2 \dot{\theta}^2$$

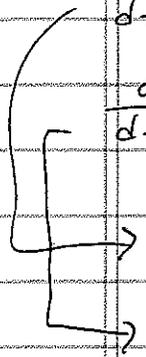
$ds - R d\theta = 0$

$$L = \frac{1}{2} M \dot{s}^2 + \frac{1}{4} M R^2 \dot{\theta}^2 + Mg s \sin \alpha$$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}} - \frac{\partial L}{\partial \theta} = -\lambda(t) R$$

$\dot{s} = R \dot{\theta}$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{s}} - \frac{\partial L}{\partial s} = \lambda(t)$$



$$\frac{1}{2} M R^2 \ddot{\theta} = -\lambda(t) R$$

$$M \ddot{s} - Mg \sin \alpha = \lambda(t) = -\frac{1}{2} M R \ddot{\theta}$$

$M R \ddot{\theta}$

$$\ddot{\theta} = \frac{2 \sin \alpha}{3 R} g$$

as  $\alpha \rightarrow 0$   
Non-sense!!

This model is highly idealized. Not enough for real rolling wheels.

$\lambda(t) = -\frac{1}{3} Mg \sin \alpha \leftarrow$  friction

$-\lambda(t) R = \frac{1}{3} Mg R \sin \alpha \leftarrow$  torque

Some ending comments (Advanced)  
(Reading Material.)

①  $L \Rightarrow \alpha L$  : doesn't change a thing.  
"Mechanical similarity"

For example,  $L = \frac{1}{2} m v^2 - mgy$  Projectile motion

Scale ~~space~~ space  $y, x \rightarrow Ay, Ax$

Scale ~~time~~ time  $t \rightarrow \sqrt{A} t$

$L \Rightarrow A L$   $v \rightarrow \sqrt{A} v$

~~scribble~~ ~~scribble~~ h  $\propto v_0^2$   
when thrown upwards

$y \propto t^2$   
 $y \propto v^2$

$L = \frac{1}{2} m \dot{x}^2 - \frac{1}{2} k x^2$

$x \rightarrow Ax$   
 $t \rightarrow t$

$L \Rightarrow A^2 L$

Period ~~time~~ is independent of amplitude!

$L = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - \frac{\alpha}{r}$  Kepler prob.

$x, y, z \rightarrow Ax, Ay, Az$

$t \rightarrow A^{3/2} t$

$L \Rightarrow \frac{1}{A} L$

$T \propto L^{3/2}$   $T$ : period  
 $L$ : orbit dimension

Kepler's third law!

Related to the Lagrangian in a non-inertial frame

②  $L \Rightarrow L + \frac{df}{dt}$  : doesn't change a thing.  
 $f = f(q, \dot{q}, t)$