

Lecture 9 Topics

- Three occupation number plots
- Classical limits
- Fermi Gas
 - Fermi energy
 - Cold metal
 - Contact potential
- Boson Gas
 - Superfluid: He₄

Boltzman Distribution/Probability

Boltzman probability:

$$P(E_n) = \frac{e^{-\frac{E_n}{k_B T}}}{\sum_n e^{-\frac{E_n}{k_B T}}}$$

Average Energy:

$$\bar{E} = \sum E_n P(E_n) = \frac{\sum_n E_n e^{-\frac{E_n}{k_B T}}}{\sum_n e^{-\frac{E_n}{k_B T}}} = \frac{\sum_n E_n \mathcal{N}(E_n)}{\sum_n \mathcal{N}(E_n)} = \frac{\int E \mathcal{N}(E) D(E) dE}{\int \mathcal{N}(E) D(E) dE}$$

Occupation number:

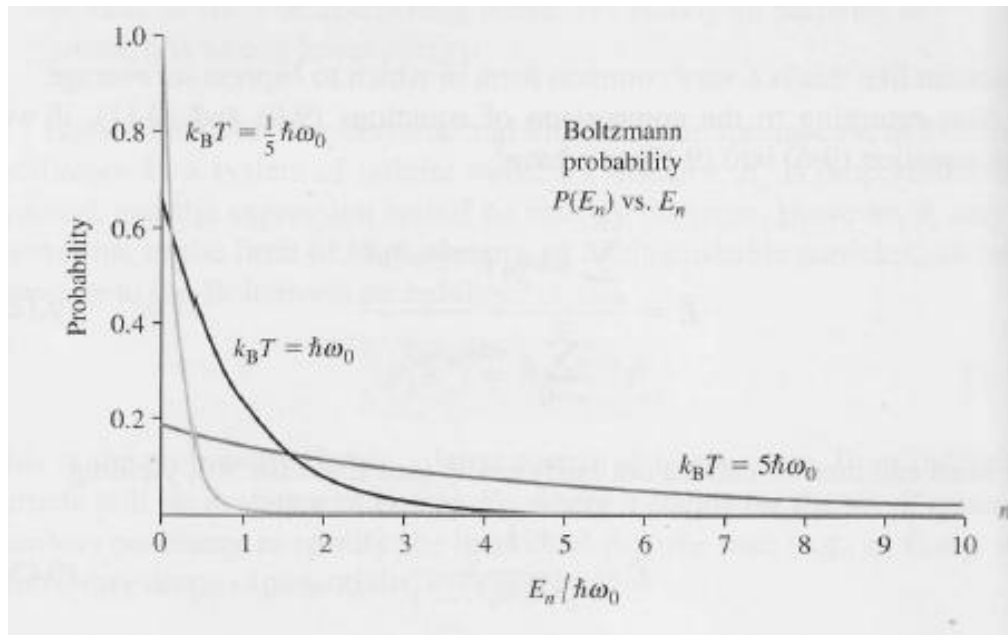
$$\mathcal{N}(E_n)_{\text{Boltzman}} = N P(E_n) = N \frac{e^{-\frac{E_n}{k_B T}}}{\sum_n e^{-\frac{E_n}{k_B T}}}$$

Density of states over (E, E+dE):

$$D(E) \equiv \frac{dn}{dE}$$

→ When E spacing is very close

Boltzmann Probability



Three types of distribution

Distribution	Occupation index	Particles	Identical particles?	Spin	Distinguishable?	Exclusion principle?
Boltzman	$\frac{1}{B e^{E/k_B T}}$	Classical	Yes	Any spin	Yes	No
Bose-Einstein	$\frac{1}{B e^{E/k_B T} - 1}$	Bosons	Yes	0 or integer spin	No	No
Fermi-Dirac	$\frac{1}{B e^{E/k_B T} + 1}$	Fermions	Yes	1/2	No	Yes

When N is REALLY large, like 10^{23}

System requirement

- Conservation of particles

$$\sum \mathcal{N}(E_i) = \mathcal{N}(E_1) + \mathcal{N}(E_2) + \cdots + \mathcal{N}(E_k) = N$$

- Conservation of energy

$$\sum \mathcal{N}(E_i) E_i = \mathcal{N}(E_1) E_1 + \mathcal{N}(E_2) E_2 + \cdots + \mathcal{N}(E_k) E_k = E$$

- Example: 4 (a, b, c, d) particles with total energy of

$$2\hbar\omega_0.$$

Maxwell-Boltzman

n	10 possible ways										No. of possibilities where a particle can have n quantum number (#)	Probability of a particle having the n quantum number ($P = \#/40$)	Probable number of particles to have the n quantum number $P \times 4$
2	<i>a</i>	<i>b</i>	<i>c</i>	<i>d</i>									
1					<i>ab</i>	<i>ac</i>	<i>ad</i>	<i>bc</i>	<i>bd</i>	<i>cd</i>			
0	<i>bcd</i>	<i>acd</i>	<i>abd</i>	<i>abc</i>	<i>cd</i>	<i>bd</i>	<i>bc</i>	<i>ad</i>	<i>ac</i>	<i>ab</i>			
	Total												

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2	a	b	c	d							4	0.1	0.4
1					ab	ac	ad	bc	bd	cd	12	0.3	1.2
0	bcd	acd	abd	abc	cd	bd	bc	ad	ac	ab	24	0.6	2.4
	Total										40	1.0	4.0

Bose-Einstein

n	10 possible ways										No. of possibilities where a particle can have n quantum number (#)	Probability of a particle having the n quantum number ($P = \#/40$)	Probable number of particles to have the n quantum number $P \times 4$
2	x	x	x	x							4	0.1	0.4
1					xx	xx	xx	xx	xx	xx	12	0.3	1.2
0	xxx	xxx	xxx	xxx	xx	xx	xx	xx	xx	xx	24	0.6	2.4
	Total										40	1.0	4.0

Bose-Einstein

n	Two possible ways		(#)	($P = \#/8$)	$P \times 4$
2	X		1	0.125	0.50
1		XX	2	0.250	1.00
0	XXX	XX	5	0.625	2.50
		Total	8	1.000	4.00

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1					xx	xx	xx	xx	xx	xx	12	0.3	1.2
0	xxx	xxx	xxx	xxx	xx	xx	xx	xx	xx	xx	24	0.6	2.4
	Total										40	1.0	4.0

Fermi-Dirac

n	1 possible way	(#)	($P=\#/8$)	$P \times 4$
2		0	0.0	0.0
1	XX	2	0.5	2.0
0	XX	2	0.5	2.0
	Total	4	1.0	4.0

Three Distributions

n	10 possible ways										No. of possibilities where a particle can have n quantum number (#)	Probability of a particle having the n quantum number ($P=\#/40$)	Probable number of particles to have the n quantum number $P \times 4$
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Probable Number Plotting

n	Maxwell-Boltzmann	Bose-Einstein	Fermi-Dirac
3	0.0	0.0	0.0
2	0.4	0.5	0.0
1	1.2	1.0	2.0
0	2.4	2.5	2.0

When there are four harmonic oscillators in a system with the system's energy is $2\hbar\omega_0$.

Probable Number



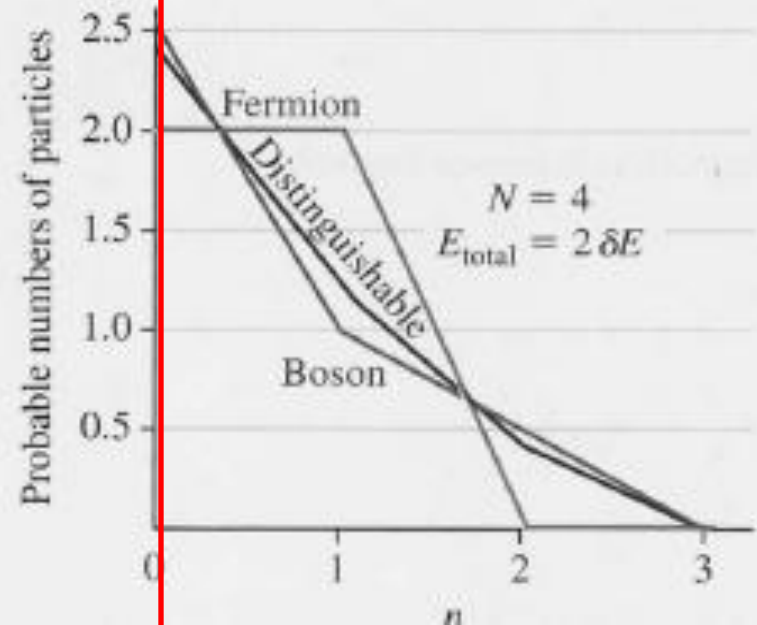
n (particle's energy state)

Probable Number Plotting

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When there are four harmonic oscillators in a system with the system's energy is $2\hbar\omega_0$.

Figure 9.9 The probable number of particles at the allowed energies depends on whether the particles are bosons, fermions, or distinguishable.

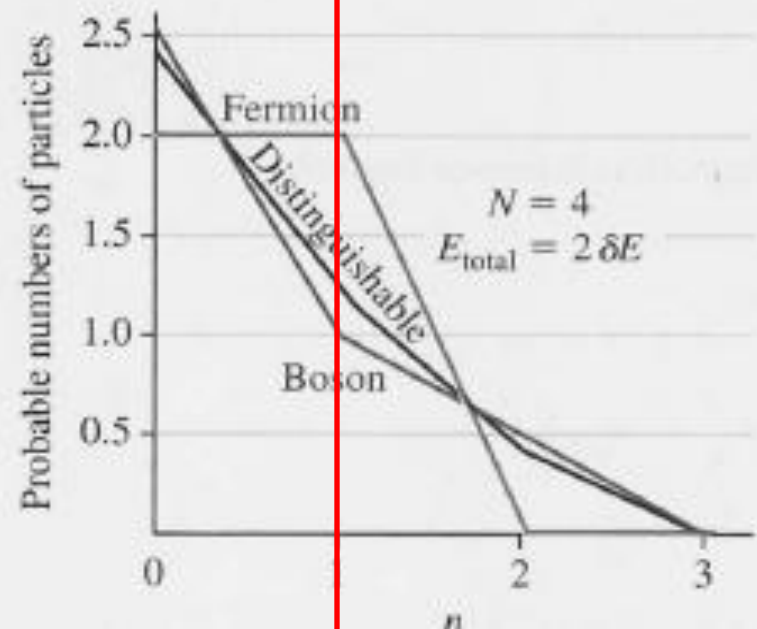


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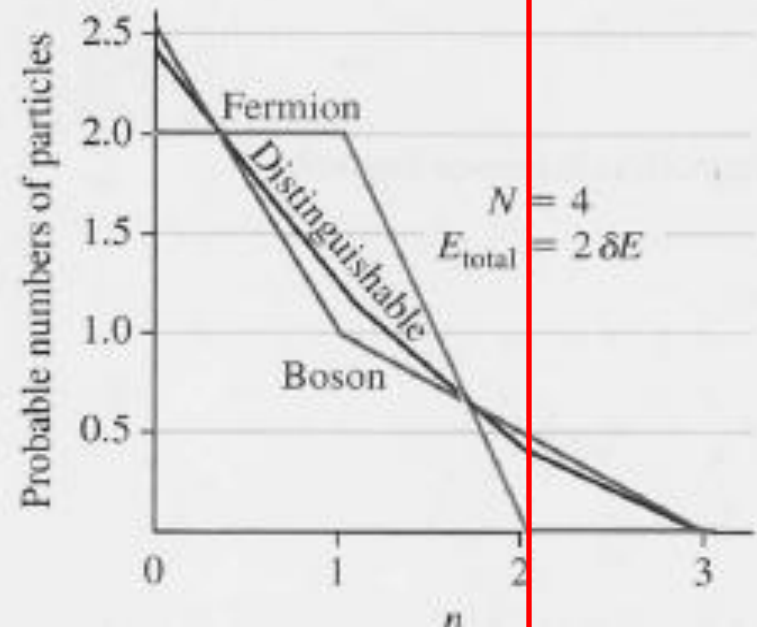


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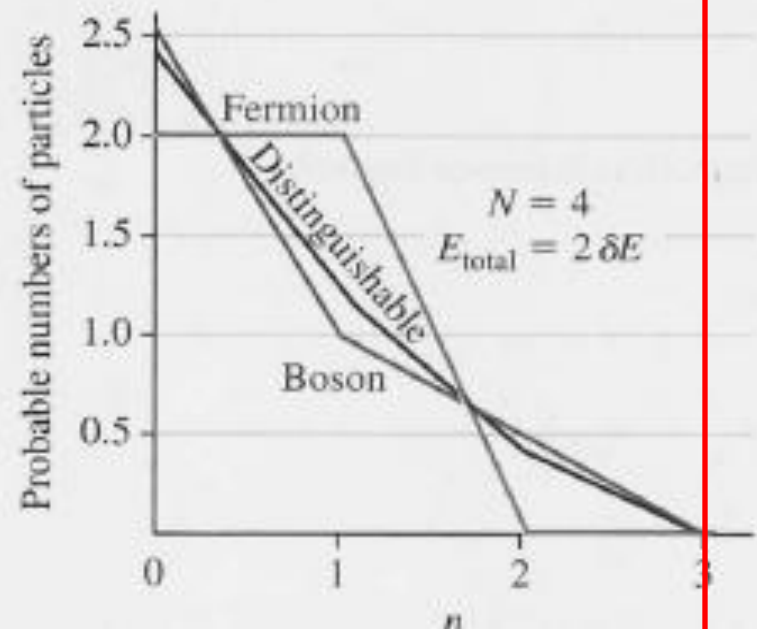


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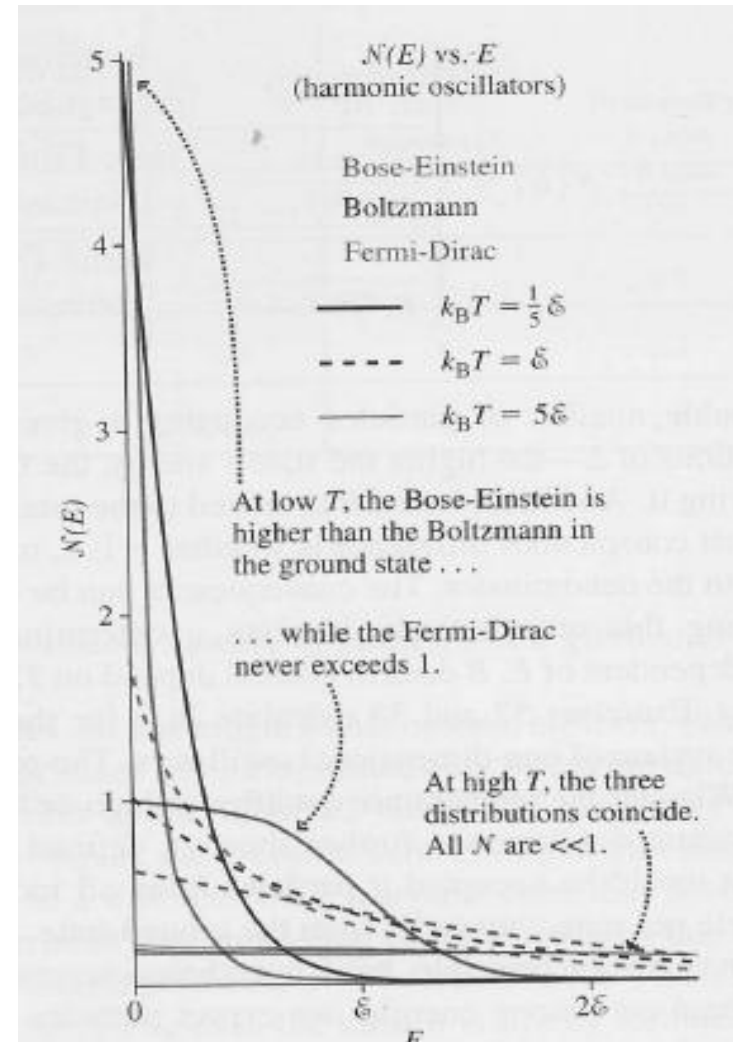
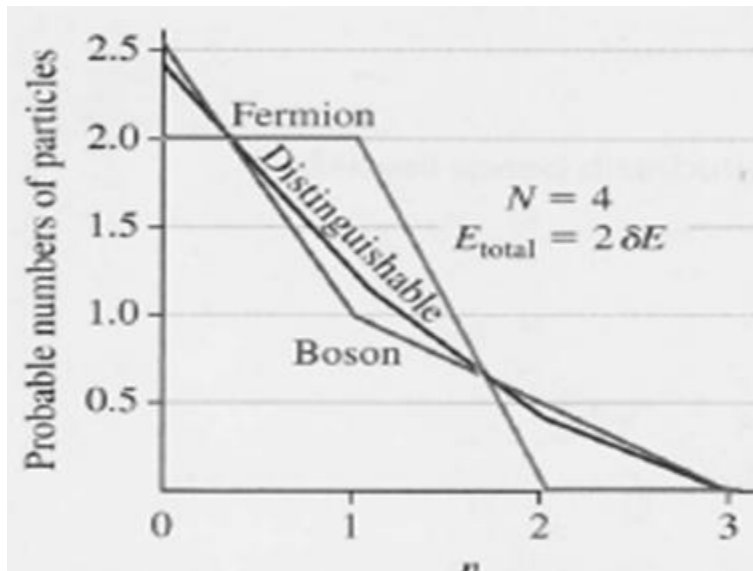


Occupation number distribution comparison

As E increases

$k_B T \ll E \rightarrow$ Low temp relative to system E

Solid lines



Occupation number distribution comparison

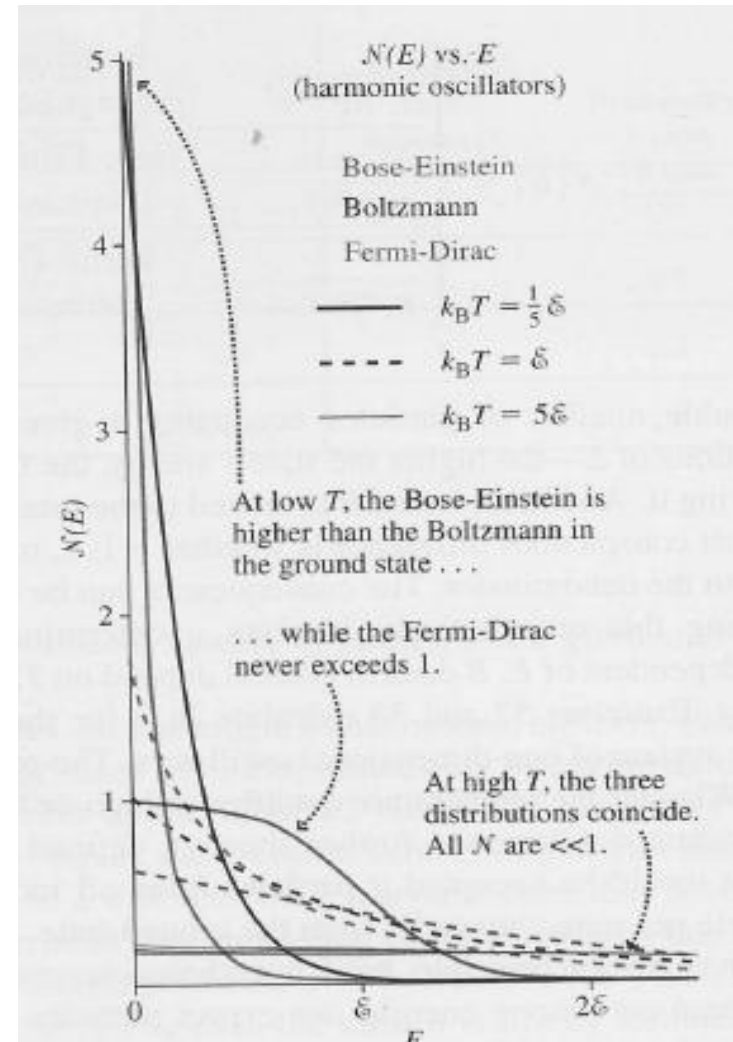
As E increases

$k_B T \ll E \rightarrow$ Low temp relative to system E

Solid lines

$k_B T \approx E$

Dashed lines



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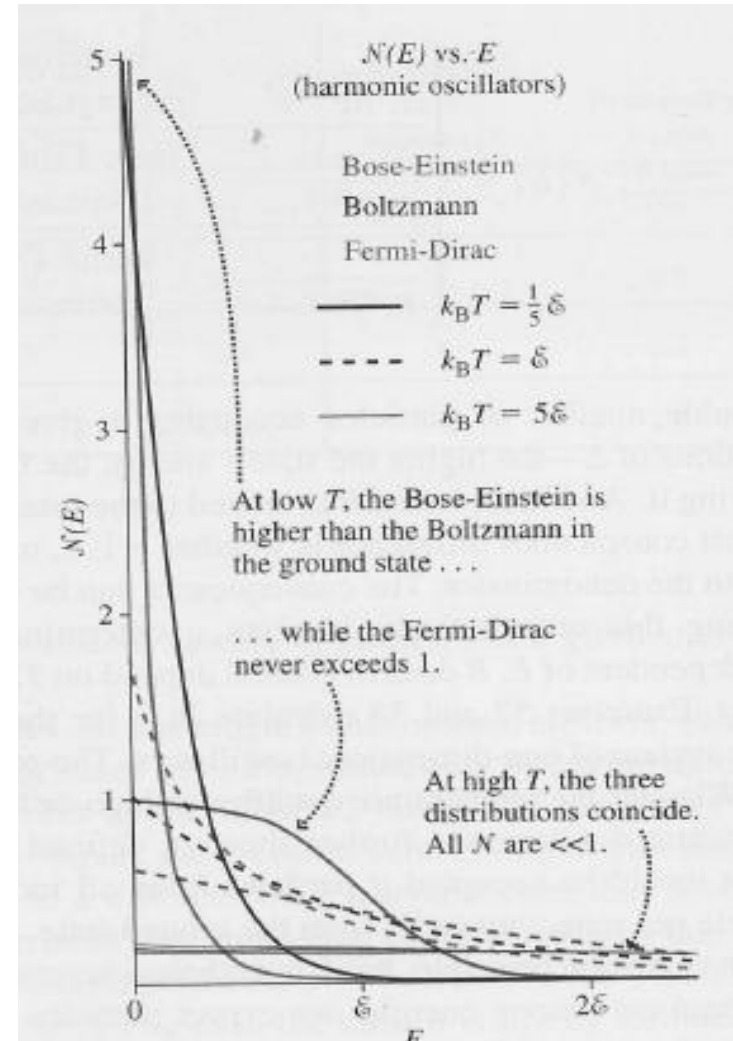
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Dashed lines

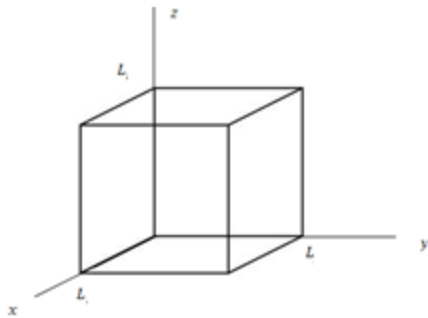
$k_B T \gg E \rightarrow$ High temp relative to system E

Thin gray lines



When to use quantum gas treatment?

- A system of N particles in a 3-D box

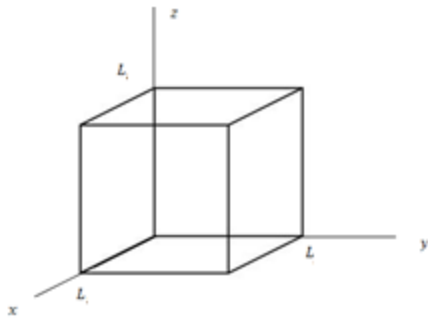


The 3D Box

$$U(\vec{x}) = \begin{cases} 0 & 0 < x, y, z < L, \\ \infty & \text{otherwise} \end{cases}$$

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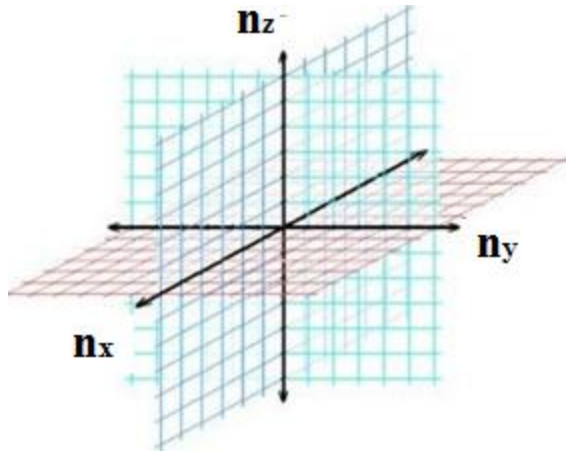
The 3D Box

$$U(\vec{x}) = \begin{cases} 0 & 0 < x, y, z < L, \\ \infty & \text{otherwise} \end{cases}$$

$$\psi(x, y, z) = A \sin \frac{n_x \pi x}{L} \sin \frac{n_y \pi y}{L} \sin \frac{n_z \pi z}{L}$$

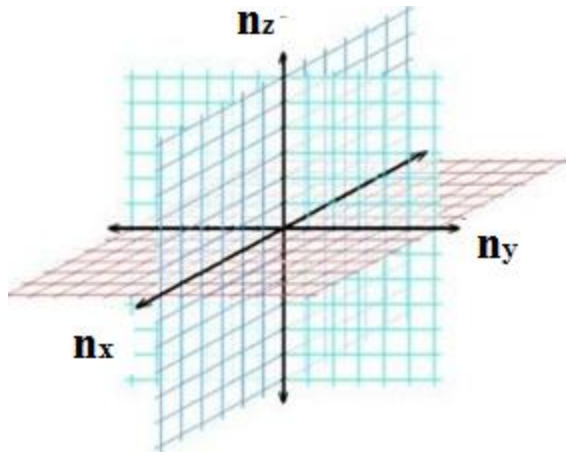
$$E_{(n_x, n_y, n_z)} = (n_x^2 + n_y^2 + n_z^2) \frac{\pi^2 \hbar^2}{2mL^2}$$

Quantum Number Space



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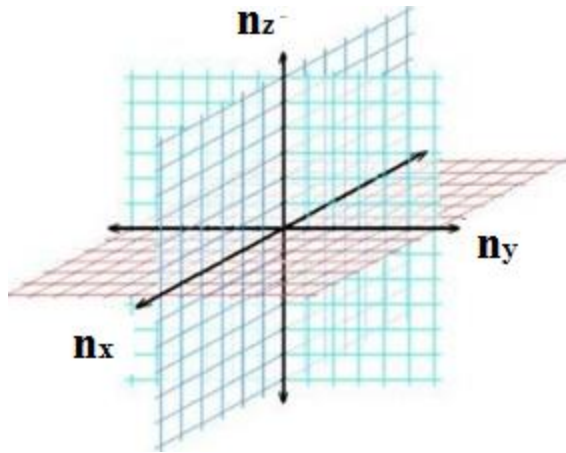
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$$E = \frac{\pi^2 \hbar^2 n^2}{2mL^2}$$

$$n = \sqrt{\frac{2mL^2}{\pi^2 \hbar^2}} \sqrt{E}$$

Quantum Number Space



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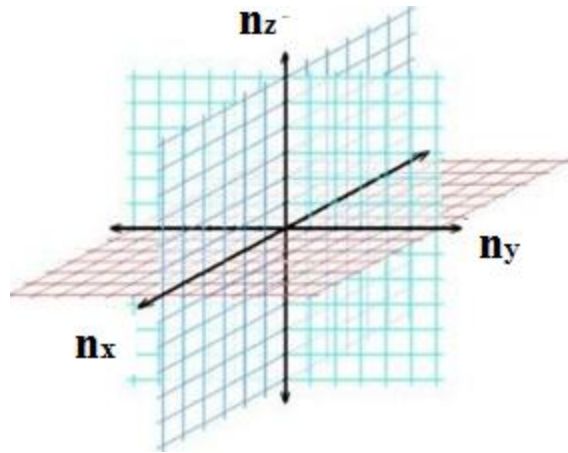
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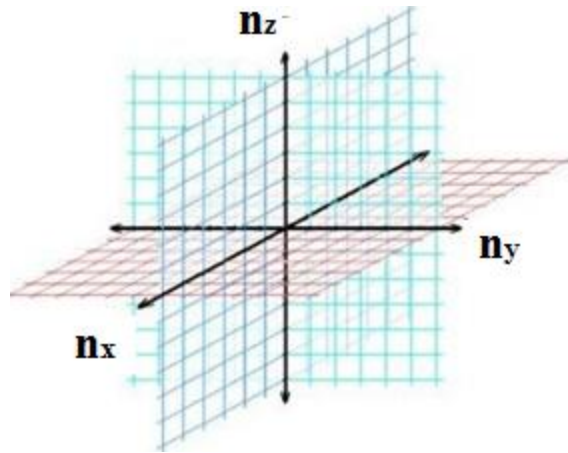
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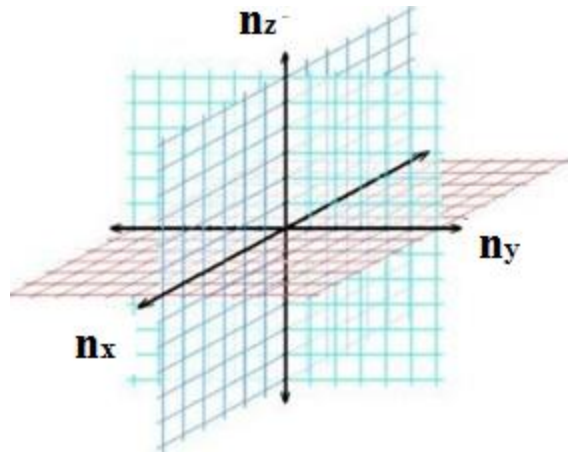
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Quantum Number Space



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Particles that obey the Boltzman distribution

$$\mathcal{N}(E) = \frac{1}{B e^{E/k_B T}}$$

Particles that obey the Bose-Einstein distribution

$$\mathcal{N}(E) = \frac{1}{B e^{E/k_B T} - 1}$$

Particles that obey the Fermi-Dirac distribution

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$$N = \int_0^{\infty} \mathcal{N}(E) D(E) dE =$$



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Classical Limit: $\frac{\hbar^3}{(m k_B T)^{3/2}} \left(\frac{N}{V}\right)^1 \ll 1$

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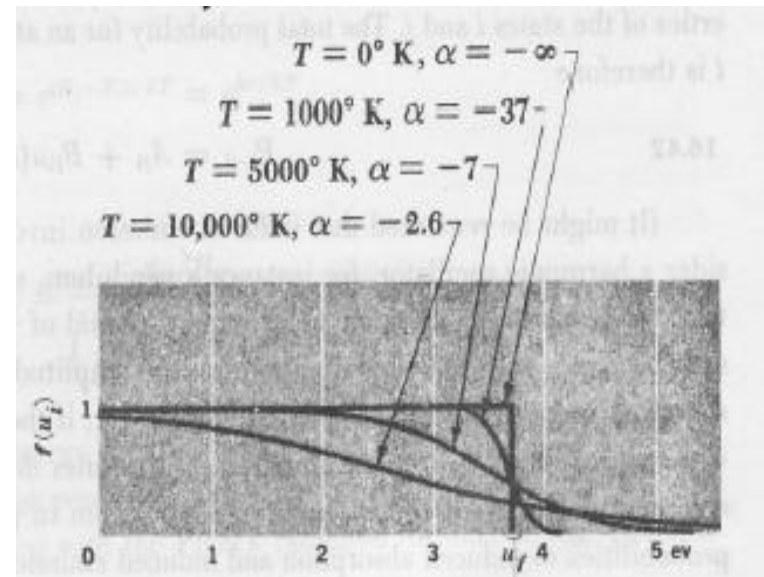
For the ideal gas, this term is
In the order of $\cong 9 \times 10^{-9}$

Classical Limit: $\frac{\hbar^3}{(m k_B T)^{3/2}} \left(\frac{N}{V}\right)^1 \ll 1$ OR $\left(\frac{\lambda}{d}\right)^3 \ll 1$

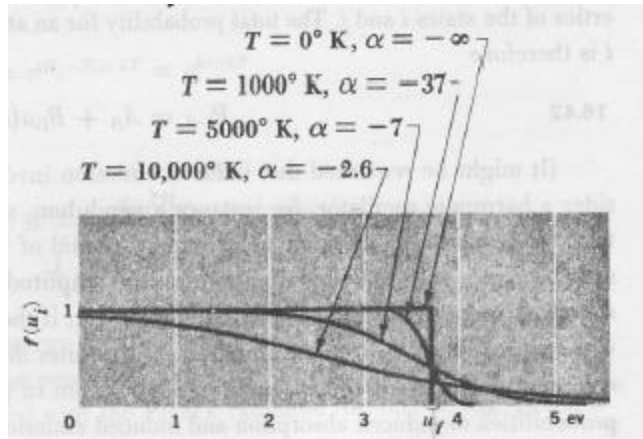
Fermi Energy

$$\mathcal{N}(E_F) = \frac{1}{B e^{E/k_B T} + 1} = \frac{1}{2} \quad \text{which makes } B = e^{-\frac{E_F}{k_B T}}$$

$$\mathcal{N}(E_F) = \frac{1}{e^{(E-E_F)/k_B T} + 1}$$



Fermi Energy



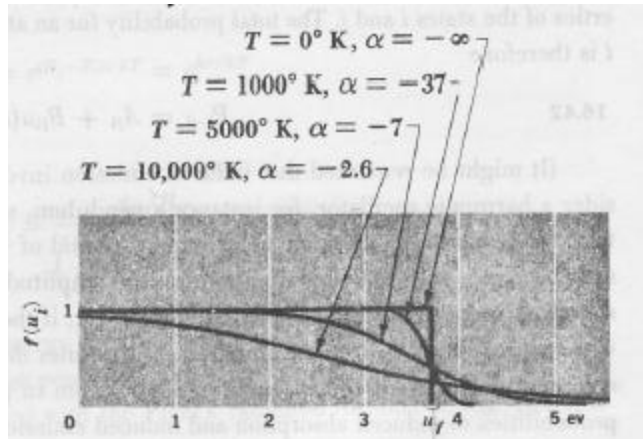
At $T=0$, electrons fill up to the E_F Level
There are N electrons.

$$N = \int \mathcal{N}(E) D(E) dE$$

$$\mathcal{N}(E) = \frac{1}{e^{(E-E_F)/k_B T} + 1}$$

$$D(E) = (2s + 1) \frac{m^{3/2} V}{\pi^3 \hbar^3 \sqrt{2}} \sqrt{E}$$

Fermi Energy



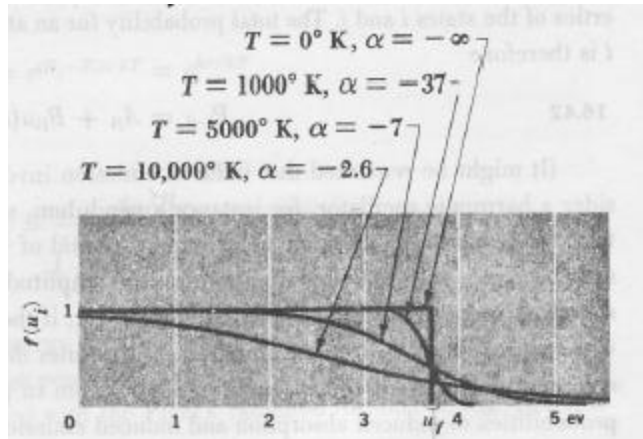
At $T=0$, electrons fill up to the E_F Level
There are N electrons.

$$N = \int \mathcal{N}(E)D(E)dE \quad N = \int_0^{E_F} \mathcal{N}(E)D(E)dE$$

$$\mathcal{N}(E) = \frac{1}{e^{(E-E_F)/k_B T} + 1} \quad =1 \text{ upto } E_F$$

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Fermi Energy



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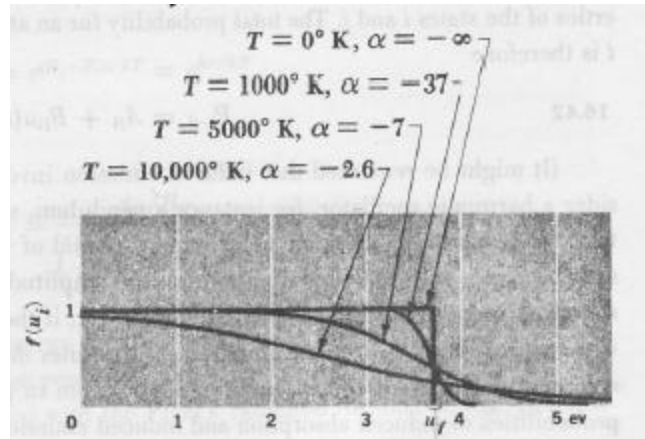
$$\mathcal{N}(E) = \frac{1}{e^{(E-E_F)/k_B T} + 1} \quad =1 \text{ upto } E_F$$

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Fermi Energy



At $T=0$, electrons fill up to the E_F Level
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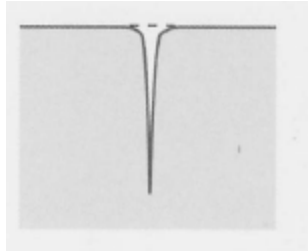
$$= 1 \text{ upto } E_F$$

$$N = \int_0^{E_F} \mathcal{N}(E) D(E) dE = \int_0^{E_F} (2s + 1) \frac{m^{3/2} V}{\pi^3 \hbar^3 \sqrt{2}} \sqrt{E} dE = (2s + 1) \frac{m^{3/2} V}{\pi^3 \hbar^3 \sqrt{2}} \left(\frac{2}{3} E_F^{3/2} \right)$$

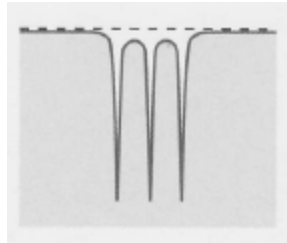
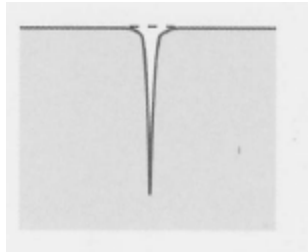
$$E_F = \frac{\pi^2 \hbar^2}{m} \left[\frac{3}{(2s + 1) \pi \sqrt{2} V} N \right]^{2/3}$$

For Silver, E_F is 5.5 eV

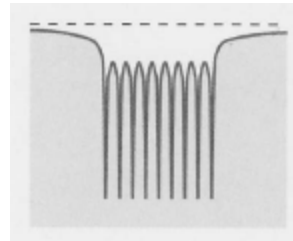
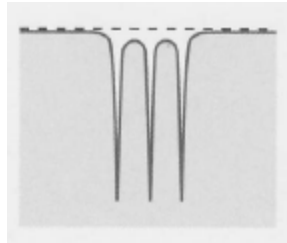
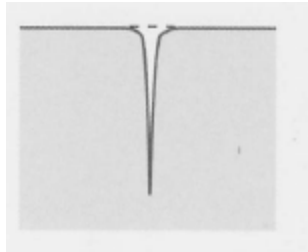
Fermi Gas: Conduction Electrons



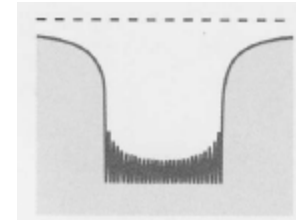
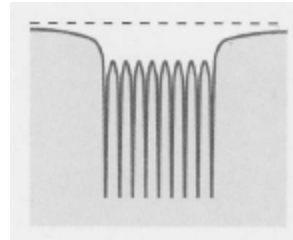
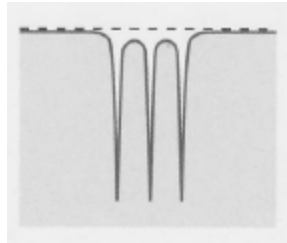
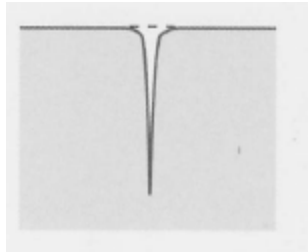
Fermi Gas: Conduction Electrons



Fermi Gas: Conduction Electrons



Fermi Gas: Conduction Electrons



Fermi Gas: Conduction Electrons

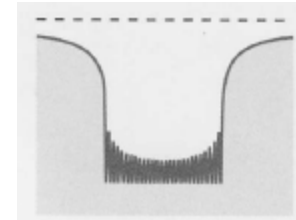
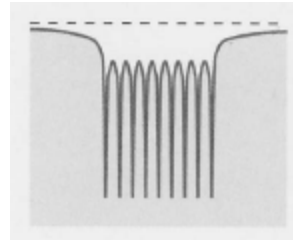
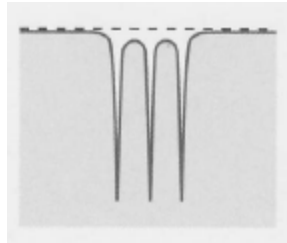
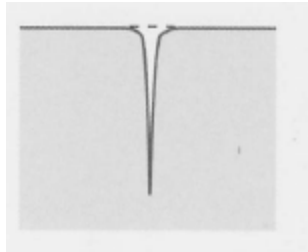
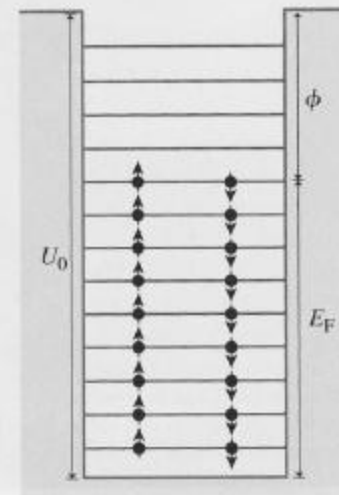
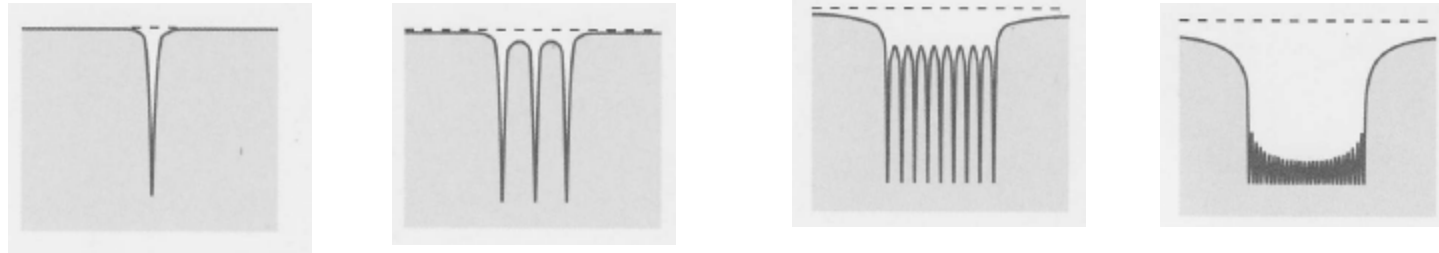


Figure 9.17 Electron energies in a "cold" metal.

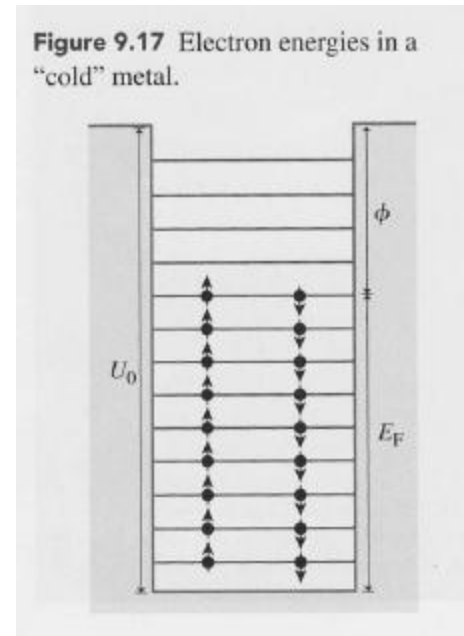


Fermi Gas: Conduction Electrons

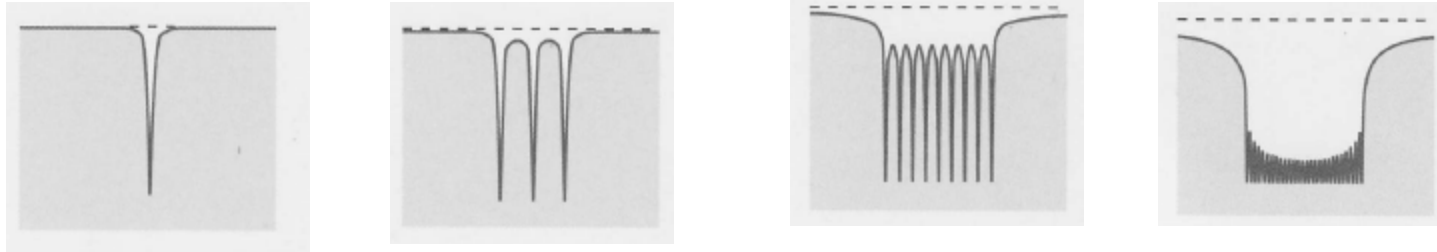


Electrons stack up the energy levels according to Pauli Exclusion Principle.

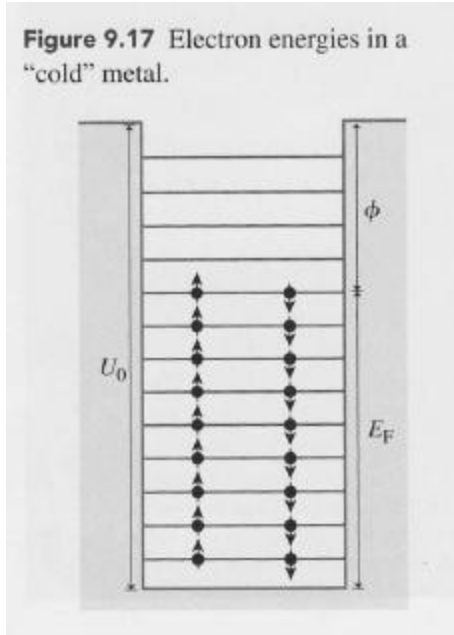
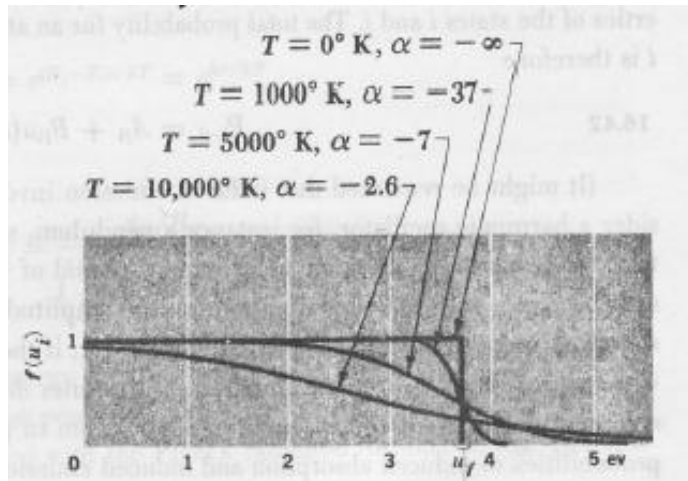
Figure 9.17 Electron energies in a "cold" metal.



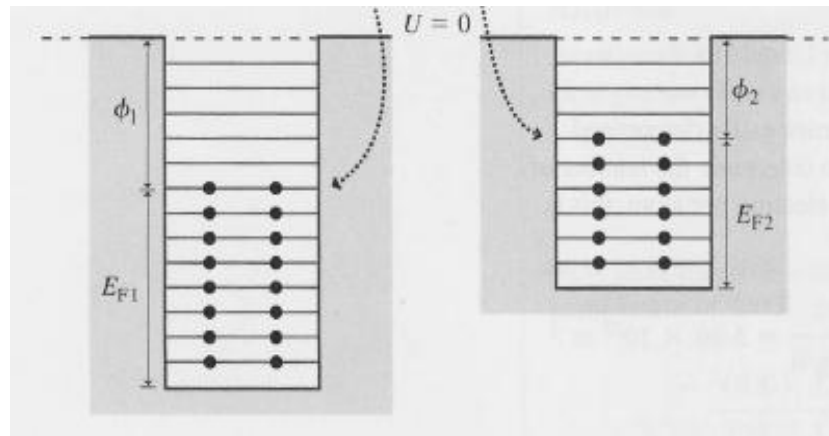
Fermi Gas: Conduction Electrons



Electrons stack up the energy levels according to Pauli Exclusion Principle.



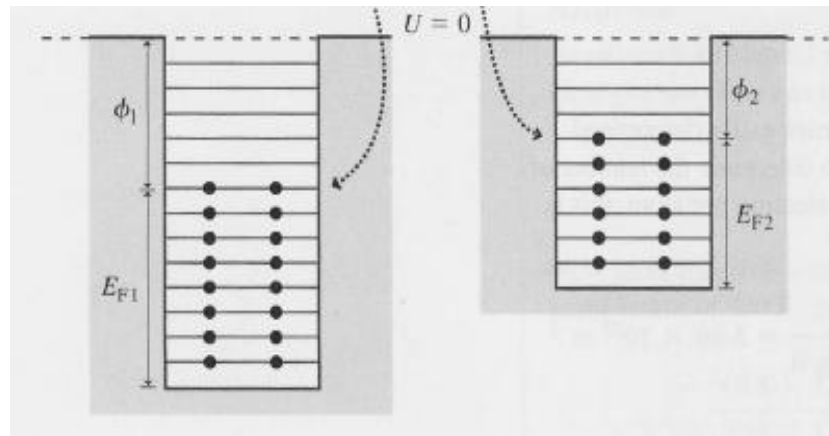
Contact Potential



Work function:

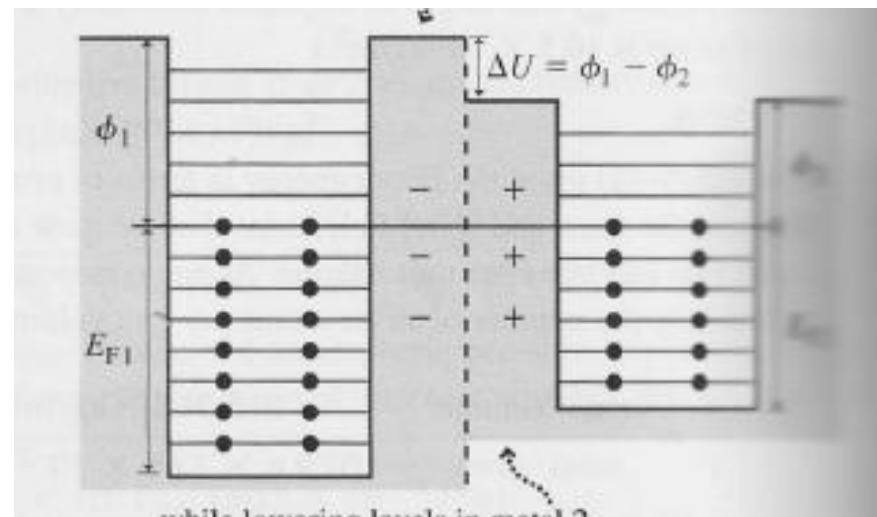
$$\phi = U_0 - E_F$$

Contact Potential

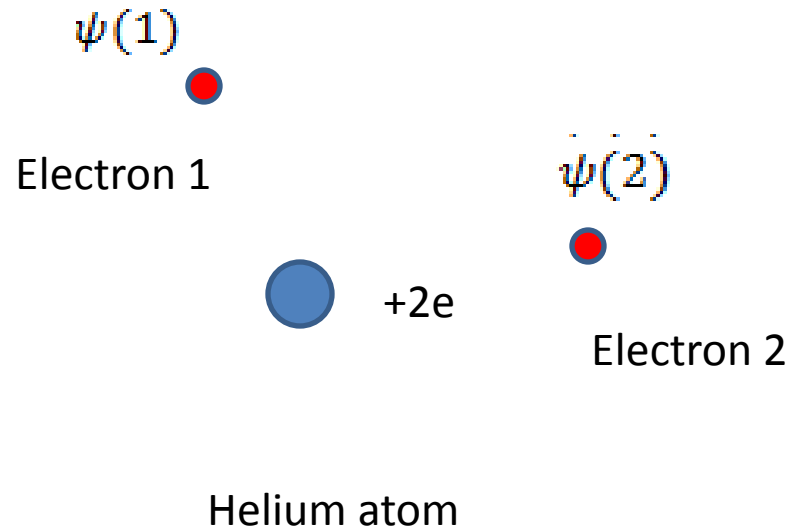


Work function:

$$\phi = U_0 - E_F$$

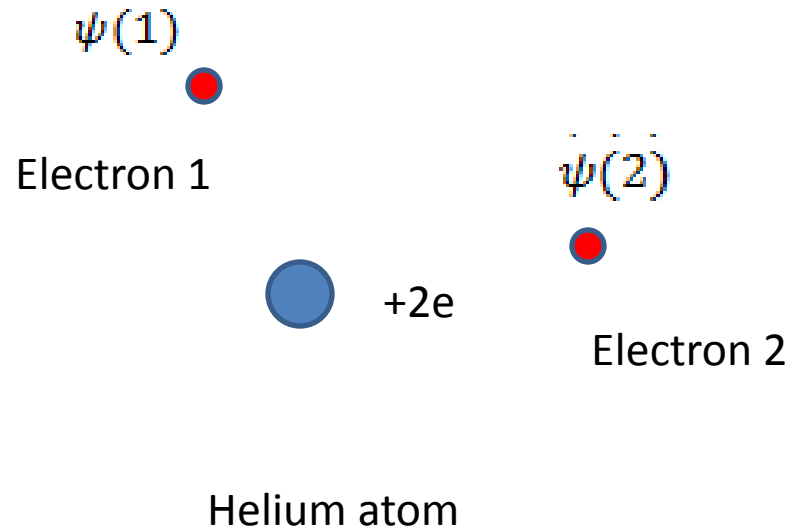


SuperFluid: He₄



SuperFluid: He₄

Is He a Boson or a Fermion?



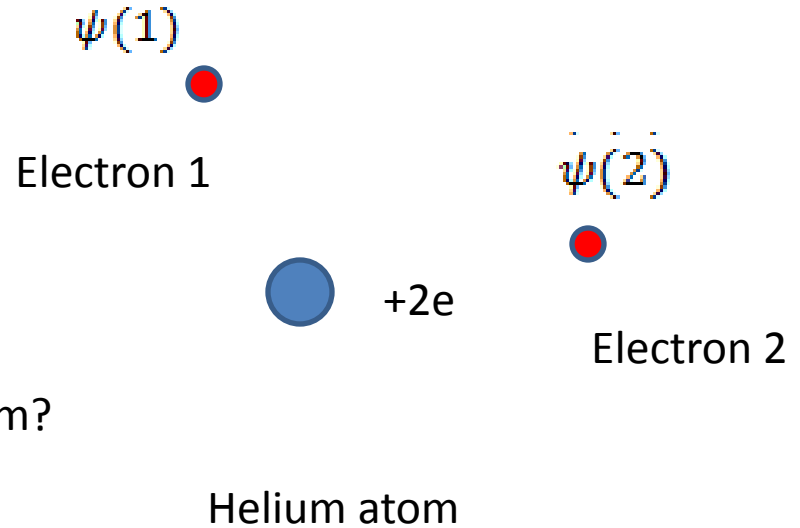
SuperFluid: He₄

He₄ has two electrons of spin 1/2

$$S_1 = \frac{1}{2} \quad S_2 = \frac{1}{2}$$

What is the total spin angular momentum?

What are m_s possibilities?



SuperFluid: He₄

He₄ has two electrons of spin 1/2

$$S_1 = \frac{1}{2} \quad S_2 = \frac{1}{2}$$

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$$S = \frac{1}{2} + \frac{1}{2} = 1 \quad \text{OR} \quad S = \frac{1}{2} - \frac{1}{2} = 0$$

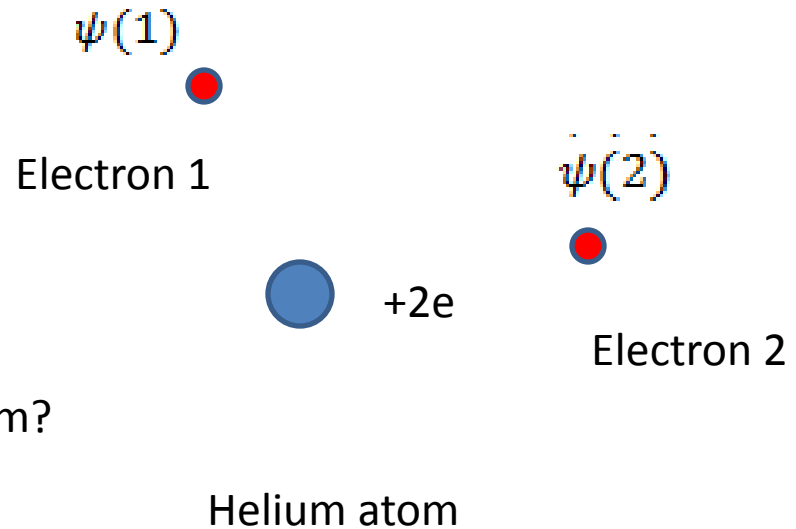
What are m_s possibilities?

For $S = 0$, $m_s = 0$

For $S = 1$, $m_s = 1$

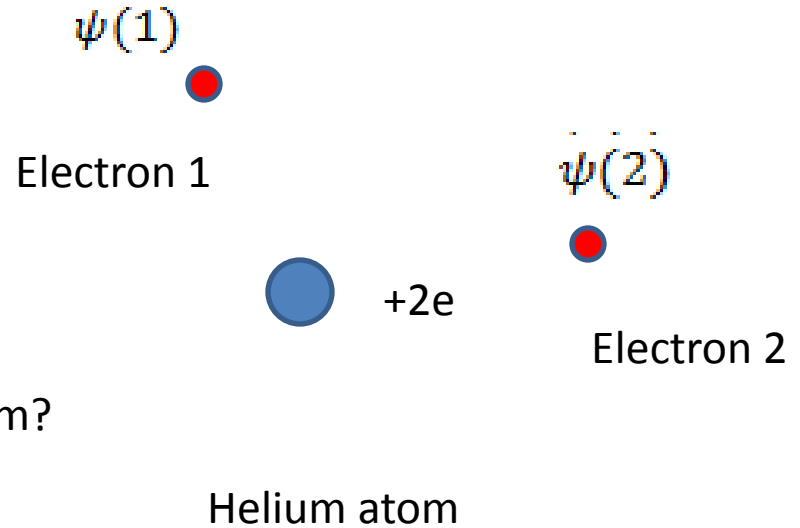
$$m_s = 0$$

$$m_s = -1$$



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$$S = \frac{1}{2} + \frac{1}{2} = 1 \quad \text{OR} \quad S = \frac{1}{2} - \frac{1}{2} = 0$$

What are m_s possibilities?

For $S = 0, m_s = 0$ $\uparrow\downarrow - \downarrow\uparrow$

For $S = 1, m_s = 1$ $\uparrow\uparrow$

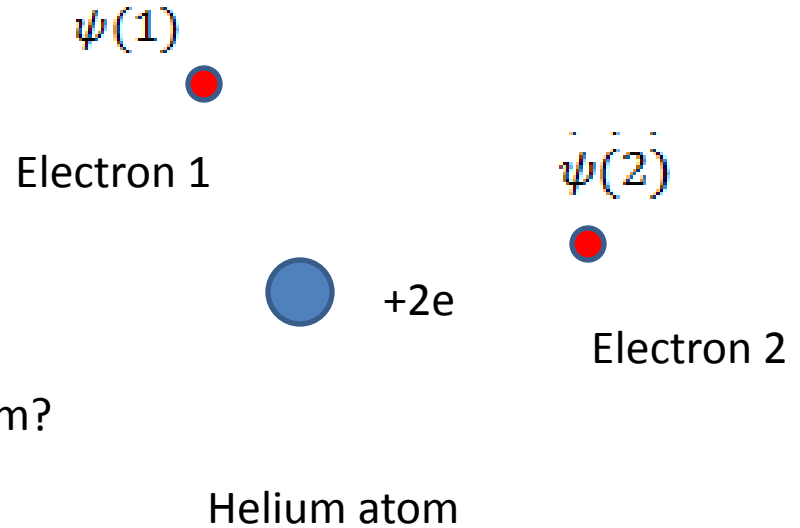
$m_s = 0$ $\uparrow\downarrow + \downarrow\uparrow$

$m_s = -1$ $\downarrow\downarrow$

SuperFluid: He₄

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What are m_s possibilities?

For $S = 0, m_s = 0$ $\uparrow\downarrow - \downarrow\uparrow$

Singlet anti-symmetric

For $S = 1, m_s = 1$ $\uparrow\uparrow$

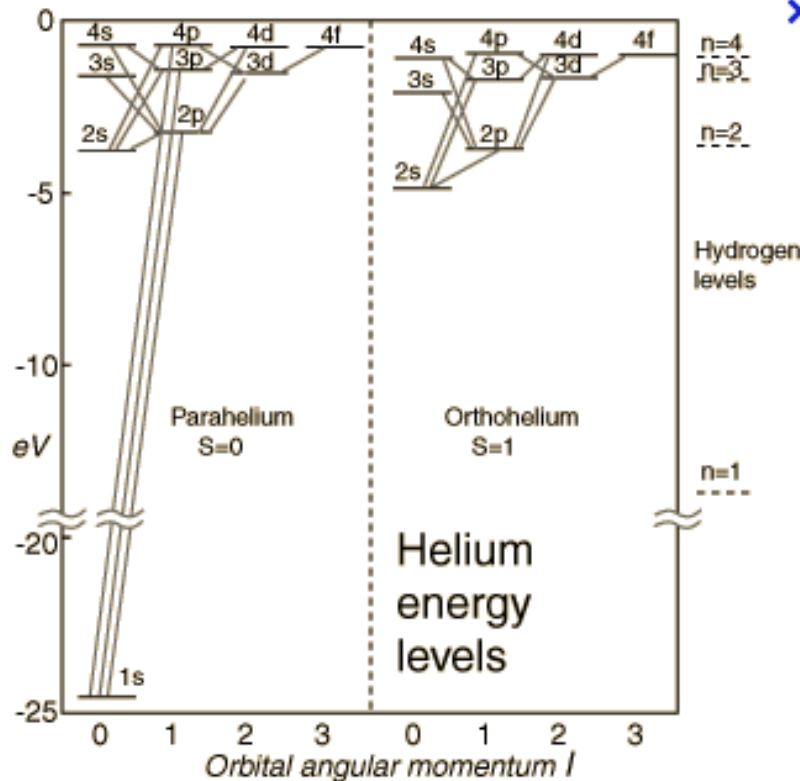
$m_s = 0$ $\uparrow\downarrow + \downarrow\uparrow$

$m_s = -1$ $\downarrow\downarrow$

Triplet symmetric



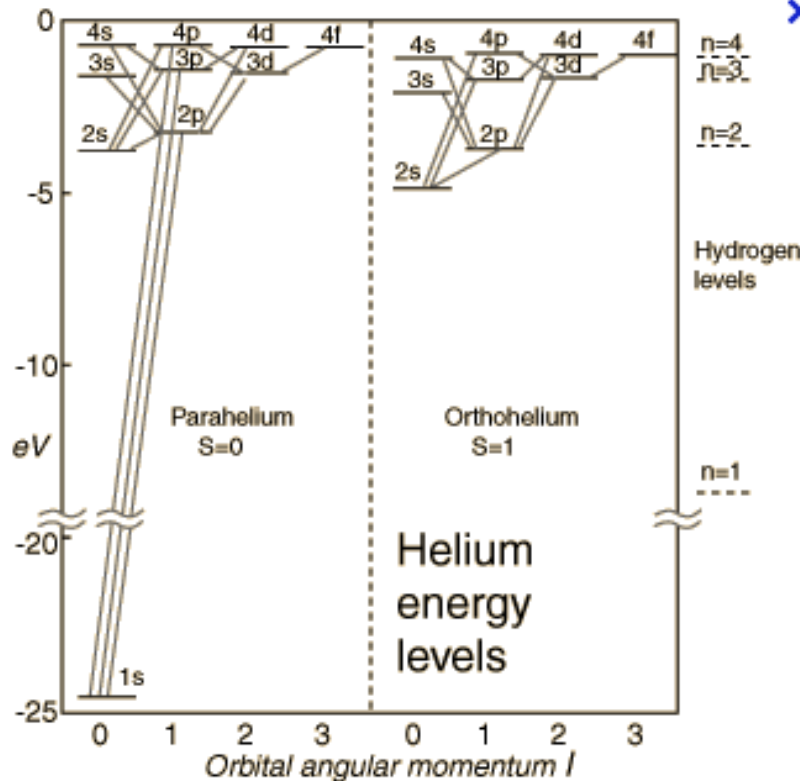
Superfluid: He4



- The Spin 0 state is more stable for the 1S orbital.
- Since electrons in He4 are considered bosons, they can occupy the same energy state.
- As He₄ cools, more stable energy states are sought by electrons.
- Around 2.2K, most electrons in He4 occupies the lowest possible energy states, meaning sharing the same wave function.

$$\frac{1}{\sqrt{2}} \left| \begin{array}{cc} \psi_{1s}(x_1) \uparrow_1 & \psi_{1s}(x_1) \downarrow_1 \\ \psi_{1s}(x_2) \uparrow_2 & \psi_{1s}(x_2) \downarrow_2 \end{array} \right|$$

Superfluid: He₄



- ✘ • Around 2.2K, most electrons in He₄ occupies the lowest possible energy states, meaning sharing the same wave function.

$$\frac{1}{\sqrt{2}} \begin{vmatrix} \psi_{1s}(x_1) \uparrow_1 & \psi_{1s}(x_1) \downarrow_1 \\ \psi_{1s}(x_2) \uparrow_2 & \psi_{1s}(x_2) \downarrow_2 \end{vmatrix}$$

- In a system of N He atoms:

$$\sum_{j=1}^N e^{-i\vec{k}_j \cdot \vec{x}_j} \begin{vmatrix} \psi_{1s}(x_1) \uparrow_1 & \psi_{1s}(x_1) \downarrow_1 \\ \psi_{1s}(x_2) \uparrow_2 & \psi_{1s}(x_2) \downarrow_2 \end{vmatrix}$$

Power of coherent waves



Power of coherent waves

